

Phys 262, HW#5: 37.29, 37.44, 37.47, 37.66, 37.69

37.29 a) AT WHAT SPEED IS REL. MOMENTUM TWICE THE CLASSICAL VALUE?

$$p = \gamma M_0 v = 2 M_0 v$$

$$\Rightarrow \gamma = 2 \Rightarrow \frac{1}{\sqrt{1-v^2/c^2}} = 2 \Rightarrow 1-v^2/c^2 = \left(\frac{1}{2}\right)^2$$

$$\Rightarrow v^2/c^2 = 1 - \left(\frac{1}{2}\right)^2 = \frac{3}{4} \Rightarrow v = \sqrt{\frac{3}{4}} c = .8660c$$

b) AT WHAT SPEED IS FORCE NEEDED TO ACCELERATE TWICE THAT NEEDED TO ACCELERATE FROM REST.

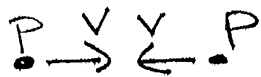
$$F = \gamma^3 M_0 a \quad \text{AT REST, } \gamma = 1 \Rightarrow \gamma^3 M_0 a = 2(M_0 a)$$

$$\Rightarrow \gamma^3 = 2 \Rightarrow \gamma = 2^{1/3} \Rightarrow \frac{1}{\sqrt{1-v^2/c^2}} = 2^{1/3} \Rightarrow 1-v^2/c^2 = \left(\frac{1}{2^{1/3}}\right)^2$$

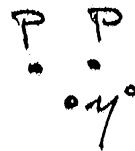
$$\Rightarrow 1-v^2/c^2 = 2^{-2/3} \Rightarrow v^2/c^2 = 1 - 2^{-2/3}$$

$$\Rightarrow v = \sqrt{1 - 2^{-2/3}} c = .608c$$

37.44



BEFORE



AFTER

REST MASS:

$$\text{PROTON: } M_p = 1.67 \times 10^{-27} \text{ kg}$$

$$\gamma: M_\gamma = 9.75 \times 10^{-28} \text{ kg}$$

a) FIND V OF PROTONS:

CONSERVATION OF ENERGY TELLS US THAT TOTAL ENERGY, E, BEFORE AND AFTER REMAINS UNCHANGED.

$$\text{MOVING} \Rightarrow E = \gamma M_0 c^2, \text{ AT REST} \Rightarrow E = M_0 c^2$$

$$\text{BEFORE: } E_{\text{TOTAL}} = E_{p_1} + E_{p_2} = \gamma_1 M_p c^2 + \gamma_2 M_p c^2$$

$$\gamma_1 = \frac{1}{\sqrt{1 - v_1^2/c^2}}, \gamma_2 = \frac{1}{\sqrt{1 - v_2^2/c^2}}. \quad v_1 = \text{Proton \#1 speed}, v_2 = \text{Proton \#2 speed}$$

$$v_1 = v_2 = v \Rightarrow \gamma_1 = \gamma_2 = \gamma = \frac{1}{\sqrt{1 - v^2/c^2}} \Rightarrow E_{\text{TOTAL}} = 2\gamma M_p c^2$$

$$\text{AFTER: } E_{\text{TOTAL}} = M_p c^2 + M_p c^2 + M_\gamma c^2 = 2M_p c^2 + M_\gamma c^2$$

$$\text{CONSERVATION} \Rightarrow 2\gamma M_p c^2 = 2M_p c^2 + M_\gamma c^2$$

$$\Rightarrow \gamma = 1 + \frac{M_\gamma}{2M_p} = 1 + \frac{9.75 \times 10^{-28}}{2(1.67 \times 10^{-27})} = 1.292$$

$$\Rightarrow \frac{1}{\sqrt{1 - v^2/c^2}} = 1.292 \Rightarrow v = \sqrt{1 - \left(\frac{1}{1.292}\right)^2} c = \sqrt{.4} c$$

$$\Rightarrow \boxed{v = .632 c}$$

b) Find K of Protons.

$$K = (\gamma - 1) m_p c^2 = (1.292 - 1)(1.67 \times 10^{-27} \text{ kg})(3 \times 10^8 \text{ m/s})^2$$

$$\Rightarrow \boxed{K = 4.4 \times 10^{-11} \text{ J}}$$

Express ANSWER IN TERMS OF MeV.

→ THIS IS KIND OF JUMPING AHEAD, BUT eV stands for electronVolts.

$$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}. \quad 1 \text{ MeV} = 1 \times 10^6 \text{ eV (one million eV)}.$$

$$4.4 \times 10^{-11} \text{ J} \times \frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ J}} = 2.74 \times 10^8 \text{ eV} = 274 \times 10^6 \text{ eV} \Rightarrow \boxed{K = 274 \text{ MeV}}$$

c) WHAT IS REST ENERGY OF γ^0 IN MeV?

$$M_{\gamma^0} c^2 = (9.75 \times 10^{-28} \text{ kg})(3 \times 10^8 \text{ m/s})^2 = 8.78 \times 10^{-11} \text{ J}$$

$$8.78 \times 10^{-11} \text{ J} \times \frac{1 \text{ eV}}{1.6 \times 10^{-19} \text{ J}} = 5.48 \times 10^8 \text{ eV} = 548 \times 10^6 \text{ eV}$$

$$\Rightarrow \boxed{M_{\gamma^0} c^2 = 548 \text{ MeV}}$$

d) NOTICE THAT $M_{\gamma^0} c^2 = 2K$! THE KINETIC ENERGY OF THE PROTONS IS CONVERTED INTO THE γ^0 'S MASS. MASS IS ENERGY.

37.47 SUN PRODUCING ENERGY AT A RATE OF 3.8×10^{26} WATT.

a) How MANY Kilograms of matter does sun lose every SECOND?

$$P = \frac{dE}{dt} \Rightarrow dE = P dt \Rightarrow \Delta E = P \Delta t = (3.8 \times 10^{26} \text{ watt})(1s) \\ = 3.8 \times 10^{26} \text{ J}$$

$$\Delta E = \Delta M c^2 \Rightarrow 3.8 \times 10^{26} \text{ J} = (\Delta M)(3 \times 10^8 \text{ m/s})^2 \\ \Rightarrow \boxed{\Delta M = 4.22 \times 10^9 \text{ kg}}$$

b) How many Tons?

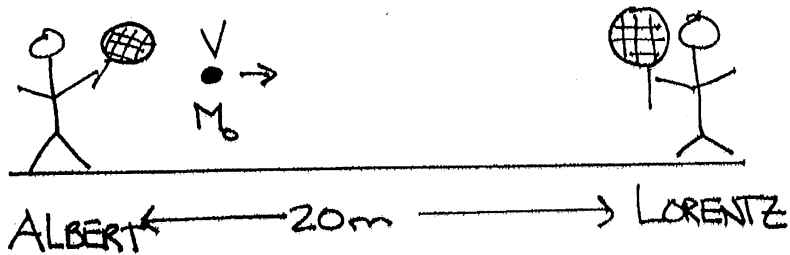
$$4.22 \times 10^9 \text{ kg} \times \frac{2.205 \text{ lb}}{\text{kg}} \times \frac{\text{ton}}{2000 \text{ lb}} = \boxed{4.655 \times 10^6 \text{ ton}} \\ = 4.655 \text{ Mton}$$

c) How long For Sun TO use up ITS MASS?

$$M_{\text{sun}} = 1.99 \times 10^{30} \text{ kg} \Rightarrow \frac{1.99 \times 10^{30} \text{ kg}}{4.22 \times 10^9 \text{ kg/s}} = 4.716 \times 10^{20} \text{ s}$$

$$4.716 \times 10^{20} \text{ s} \times \frac{\text{h}}{3600 \text{ s}} \times \frac{\text{day}}{24 \text{ h}} \times \frac{\text{year}}{365 \text{ day}} = \boxed{1.5 \times 10^{13} \text{ years}}$$

37.66



$$M_0 = .058 \text{ Kg}$$

a) IF $V = 80 \text{ m/s}$, WHAT IS KINETIC ENERGY?

AT SUCH A SMALL SPEED, WE CAN USE $K = \frac{1}{2} M V^2 = \frac{1}{2} (.058 \text{ Kg})(80 \text{ m/s})^2$

$$\Rightarrow \boxed{K = 185.6 \text{ J}}$$

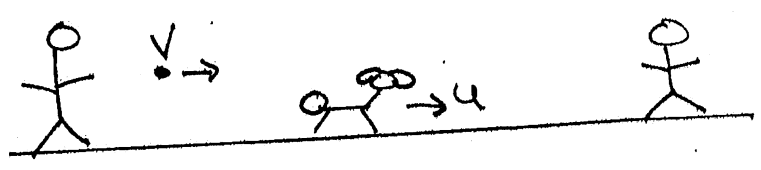
b) IF $V = 1.8 \times 10^8 \text{ m/s}$, WHAT IS KINETIC ENERGY?

HERE, WE NEED $K = (\gamma - 1) M_0 c^2$, $\gamma = \frac{1}{\sqrt{1 - (\frac{1.8}{3})^2}} = 1.25$

$\hookrightarrow \frac{1.8 \times 10^8}{3 \times 10^8} = \frac{1.8}{3}$

$$\Rightarrow K = .25 (.058 \text{ Kg})(3 \times 10^8 \text{ m/s})^2 \Rightarrow \boxed{K = 1.3 \times 10^{15} \text{ J}}$$

2) RABBIT RUNS FROM ALBERT TO LORENTZ AS BALL GOES FROM ALBERT TO LORENTZ.



LET S BE ALBERT/LORENTZ FRAME.
 S' BE RABBIT'S FRAME
 $\Rightarrow u = 2.2 \times 10^8 \text{ m/s} \rightarrow$ RABBIT'S SPEED

ALBERT AND LORENTZ SEE $V = 1.8 \times 10^8 \text{ m/s} \rightarrow$ SAME DIRECTION AS RABBIT.

CALL RABBIT'S DIRECTION X $\Rightarrow V_x = 1.8 \times 10^8 \text{ m/s}$, $u = 2.2 \times 10^8 \text{ m/s}$, $V'_x = ?$

$$V'_x = \frac{V_x - u}{1 - \frac{u V_x}{c^2}} = \frac{1.8 \times 10^8 \text{ m/s} - 2.2 \times 10^8 \text{ m/s}}{1 - \frac{(2.2)(1.8)}{3^2}} \Rightarrow \boxed{V'_x = -7.14 \times 10^7 \text{ m/s}}$$

d) WHAT DISTANCE DOES RABBIT MEASURE FROM ALBERT TO LORENTZ?

A METER STICK FROM ALBERT TO LORENTZ DOES NOT MOVE IN THEIR FRAME OF REFERENCE (S) \Rightarrow THEY MEASURE PROPER LENGTH. $\rightarrow L_0 = 20\text{m}$

RABBIT'S MEASURED LENGTH: $L = \frac{L_0}{\gamma_R}$ ($\gamma_R = \gamma$ FOR RABBIT'S SPEED)

$$\gamma_R = \frac{1}{\sqrt{1 - \left(\frac{2.2}{3}\right)^2}} = 1.47 \Rightarrow \boxed{L = \frac{20\text{m}}{1.47} = 13.6\text{m}}$$

e) HOW LONG DOES IT TAKE RABBIT TO RUN ACCORDING TO PLAYERS? RABBIT MEASURES PROPER TIME.*

PLAYERS MEASURE A ^{PROPER} DISTANCE L_0

$$\Rightarrow \Delta t = \frac{L_0}{u} = \frac{20\text{m}}{2.2 \times 10^8 \text{m/s}} \Rightarrow \boxed{\Delta t = 9.09 \times 10^{-8} \text{s}}$$

$$f) \Delta t = \gamma \Delta t_0 \Rightarrow \Delta t_0 = \frac{\Delta t}{\gamma_R} = \frac{9.09 \times 10^{-8} \text{s}}{1.47} \Rightarrow \boxed{\Delta t_0 = 6.18 \times 10^{-8} \text{s}}$$

$$\text{OR WE COULD USE } \Delta t_0 = \frac{L}{u} = \frac{13.6\text{m}}{2.2 \times 10^8 \text{m/s}} = 6.18 \times 10^{-8} \text{s}$$

(* BECAUSE THE EVENT IS THE RABBIT GOING FROM ALBERT TO LORENTZ.)

37.69

V_{rocket}	t_{Earth}	t_{rocket}	$E (J)$	$E (\%)$
.5c	1000 yr	866 yr	1.04×10^{20}	10.40%
.99c	505 yr	71.2 yr	6.38×10^{20}	6380%
.9999c	500 yr	7.07 yr	6.36×10^{21}	63641% → THAT'S A LOT

EVENT = Rocket travels to Betelgeuse $\Rightarrow t_{\text{rocket}} = \Delta t_0$.

$$\Rightarrow t_{\text{Earth}} = \Delta t \Rightarrow t_{\text{Earth}} = \gamma t_{\text{rocket}} \Rightarrow t_{\text{rocket}} = \frac{1}{\gamma} t_{\text{Earth}}$$

Earth MEASURES $L_0 \Rightarrow t_{\text{Earth}} = \frac{500 \text{ light-years}}{v}$

Moving Rocket $\Rightarrow E = \gamma M_0 c^2 = \gamma (1000 \text{ kg}) c^2 = \gamma (9 \times 10^{19} \text{ J})$

a) $V = .5c \Rightarrow \gamma = \frac{1}{\sqrt{1-.5^2}} = 1.1547, t_{\text{Earth}} = \frac{500 \text{ yr}}{.5 (\text{ly/yr})} = 1000 \text{ year}$

$\Rightarrow t_{\text{rocket}} = \frac{1000 \text{ yr}}{1.1547} = 866 \text{ yr}, E = 1.04 \times 10^{20} \text{ J}$

$\% = \frac{1.04 \times 10^{20} \times 100}{1 \times 10^{19}} = 10.4 \times 100 = 1040\%$

b) $V = .99c \Rightarrow t_{\text{Earth}} = \frac{500}{.99} = 505, \gamma = \frac{1}{\sqrt{1-.99^2}} = 7.088, \dots$ (see table)

c) $V = .9999c, \gamma = 70.712$